



The Deep Space Network

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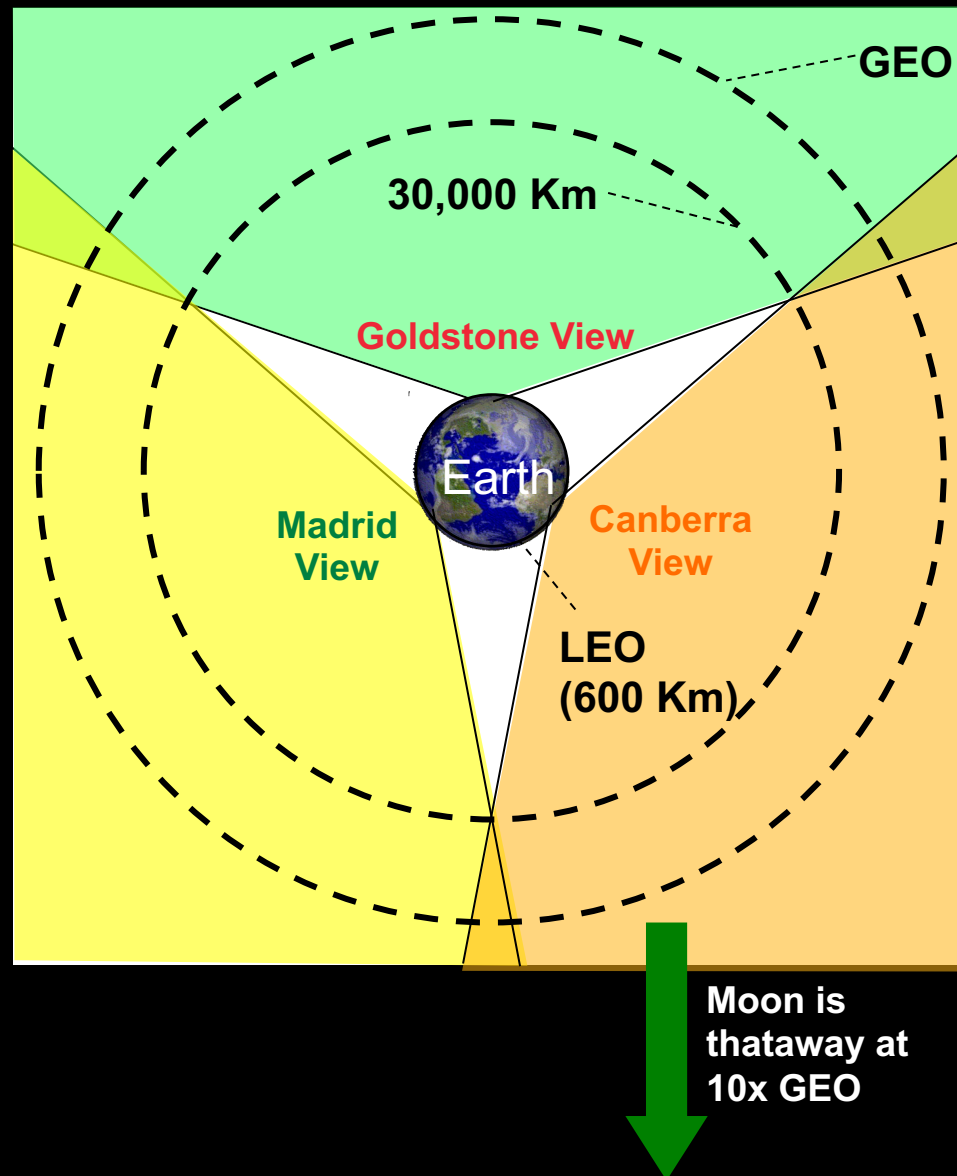


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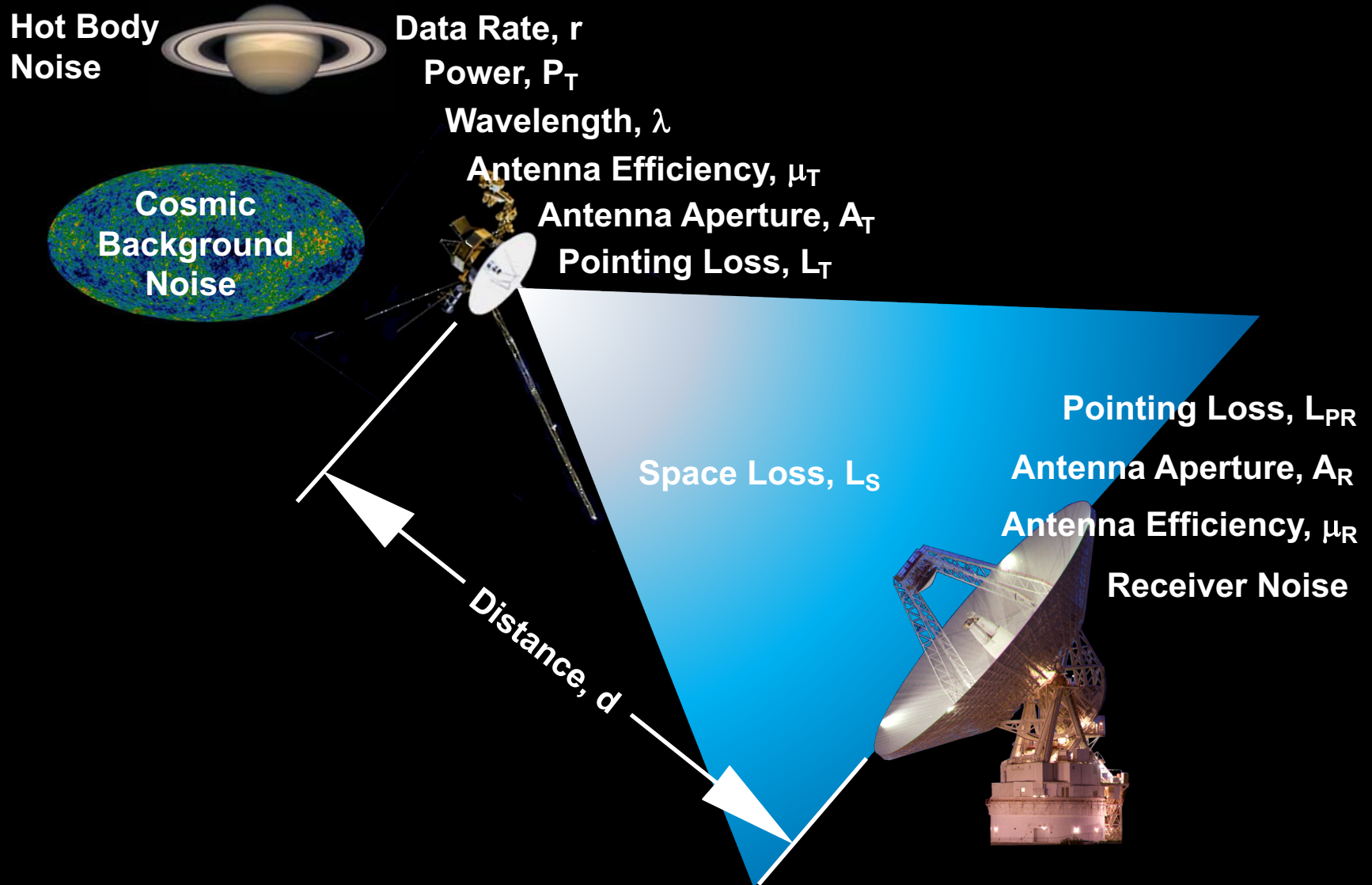
DSN Antennas in Canberra, Australia



A Global Enterprise by Necessity



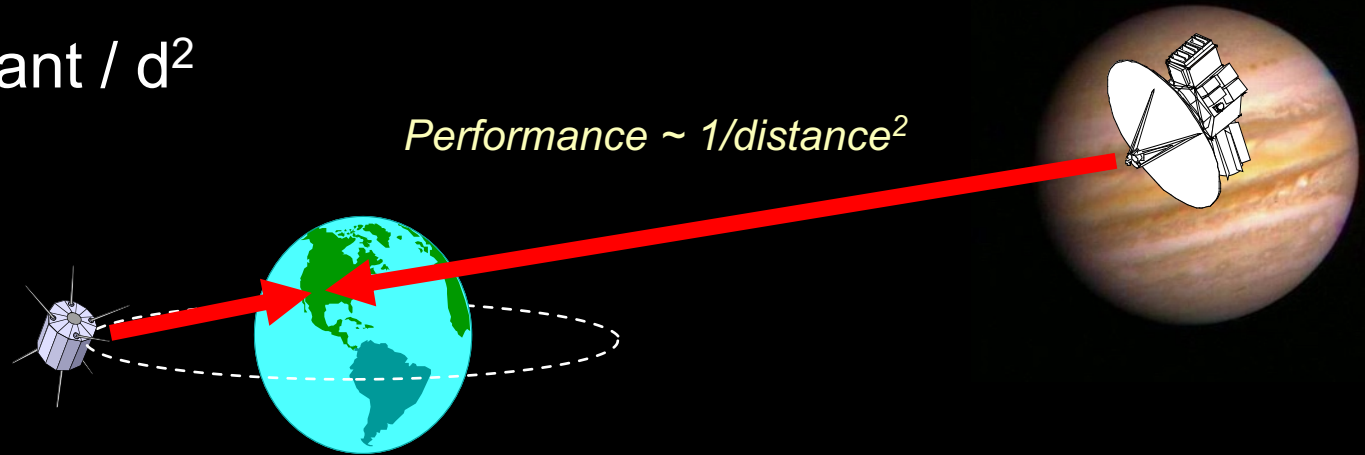
Deep Space Communications



Why Deep Space Communications is Hard

$$E_b/N_0 = \text{constant} / d^2$$

Performance ~ 1/distance²



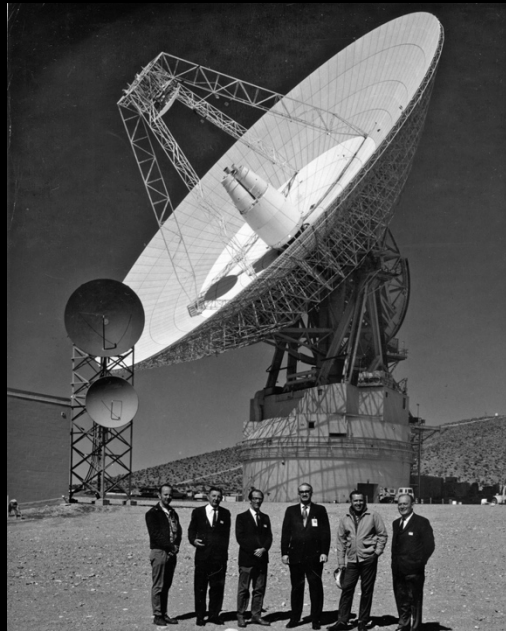
Relative Difficulty

<i>Place</i>	<i>Distance</i>	<i>Difficulty</i>
<i>GEO</i>	$4 \times 10^4 \text{ km}$	<i>Baseline</i>
<i>Moon</i>	$4 \times 10^5 \text{ km}$	<i>100</i>
<i>Mars</i>	$3 \times 10^8 \text{ km}$	5.6×10^7
<i>Jupiter</i>	$8 \times 10^8 \text{ km}$	4.0×10^8
<i>Pluto</i>	$5 \times 10^9 \text{ km}$	1.6×10^{10}

History of Ground Antennas



1958, 26m Station



1966, 64m Station



1979, 34m Station

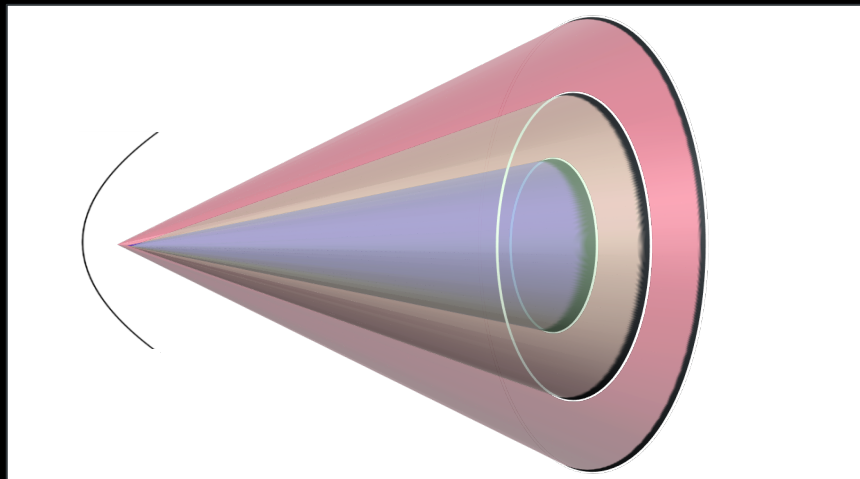


**1988, 70m Station
(converted from prior 64 antennas)**

Higher Frequency is Good

$$E_b/N_0 = \text{constant} * f^2$$

- First deep space missions transmitted at 960 MHz
- 2.2 GHz (S-band) became standard in 1969
- 8.4 GHz (X-band) became prevalent in the early 1970s
- 32 GHz (Ka-band) is now becoming the standard



Lowering the System Noise

$$E_b/N_0 = \text{constant}/T$$

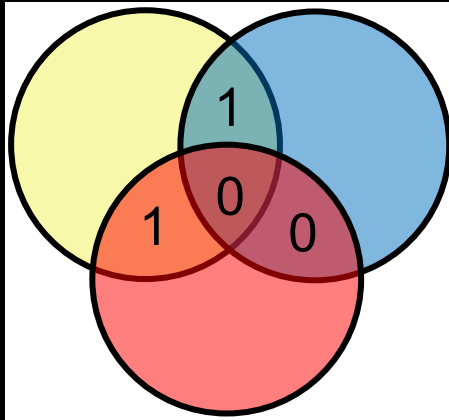
- Some of T cannot be controlled
- Focus on contributions from spacecraft and DSN
- Avoid interference
 - Our own spectrum from the ITU
- Best low noise amplifiers we can
 - Physical temperature is ~ 12 K



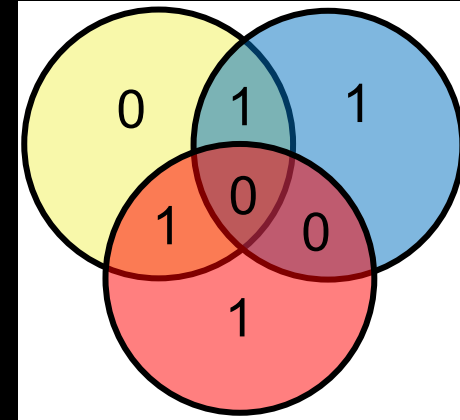
Ka-band (32 GHz) low noise amplifier

Error Korecting Cods

An example of coding: the (7, 4) Hamming code

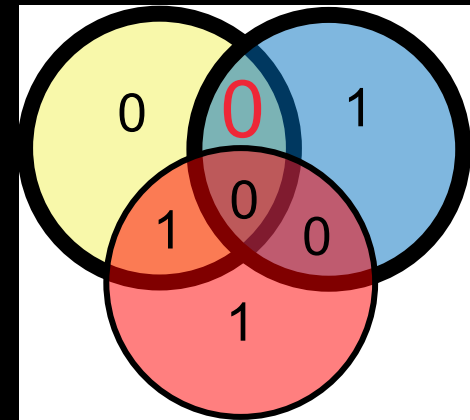
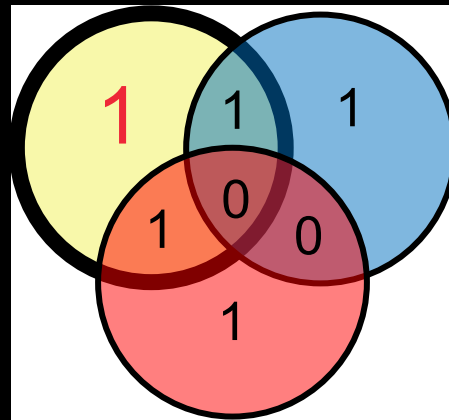


Place 4 information bits in the intersections of the Venn diagram



Fill in the diagram so that the circles have an even number of 1's

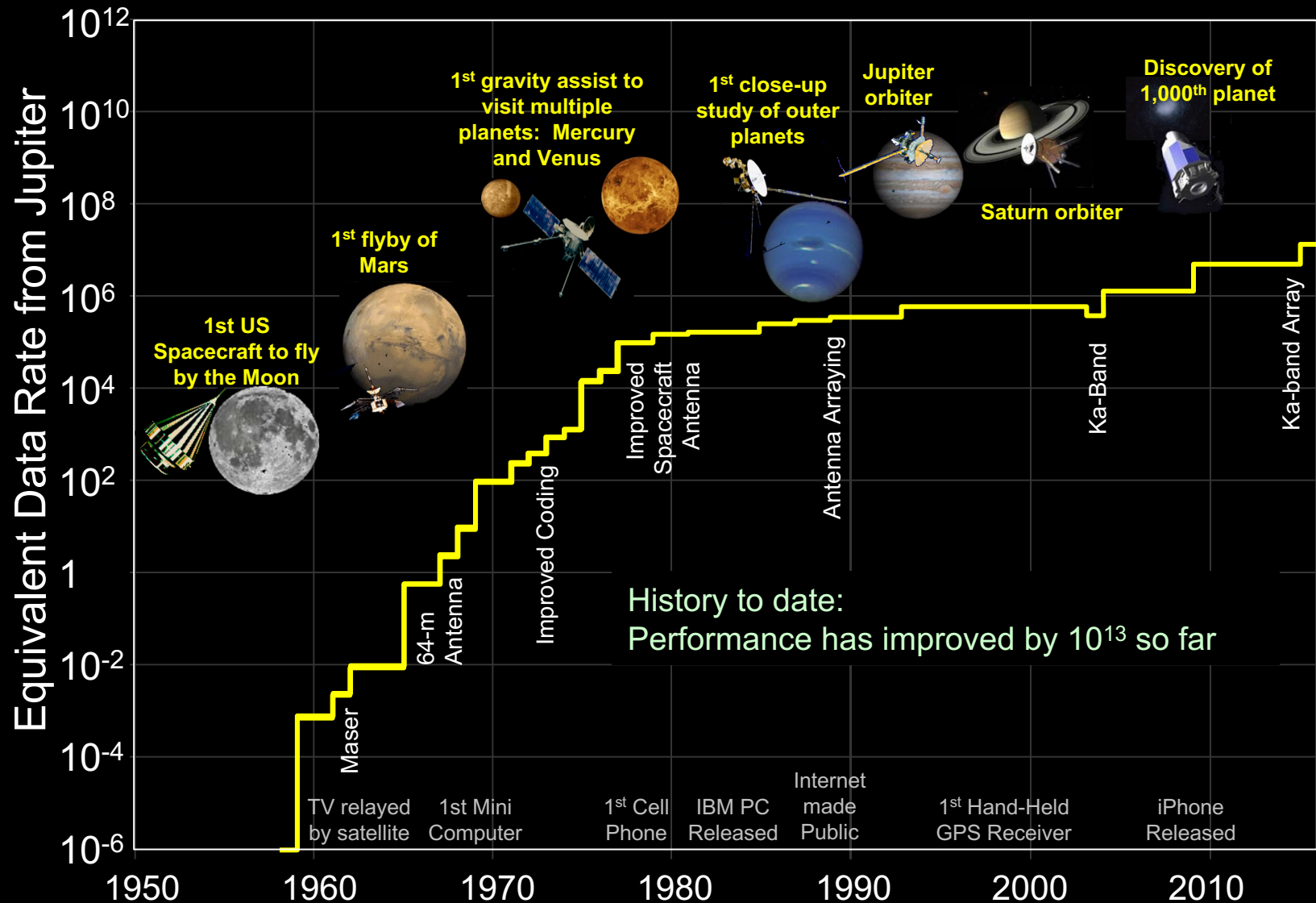
If a single error occurs, it can be corrected by locating the circles with an odd number of 1's and changing the bit in their intersection



Compression – Being stingy with bits

- Data compression is like texting
 - FYI JOE'L BRB 2 HELP L8R
 - Four your information, Joe will be right back to help later
 - Compression ratio = 39:24, or almost 2:1
- Images can be compressed 10:1
- Videos and hyperspectral images even more
- Even better: Use data onboard to answer questions and only send the answers!
 - Navigation – where am I now?
 - Locating interesting areas in a scene
 - Onboard science

A History of Improving Communications



Navigation using the Communications Signal

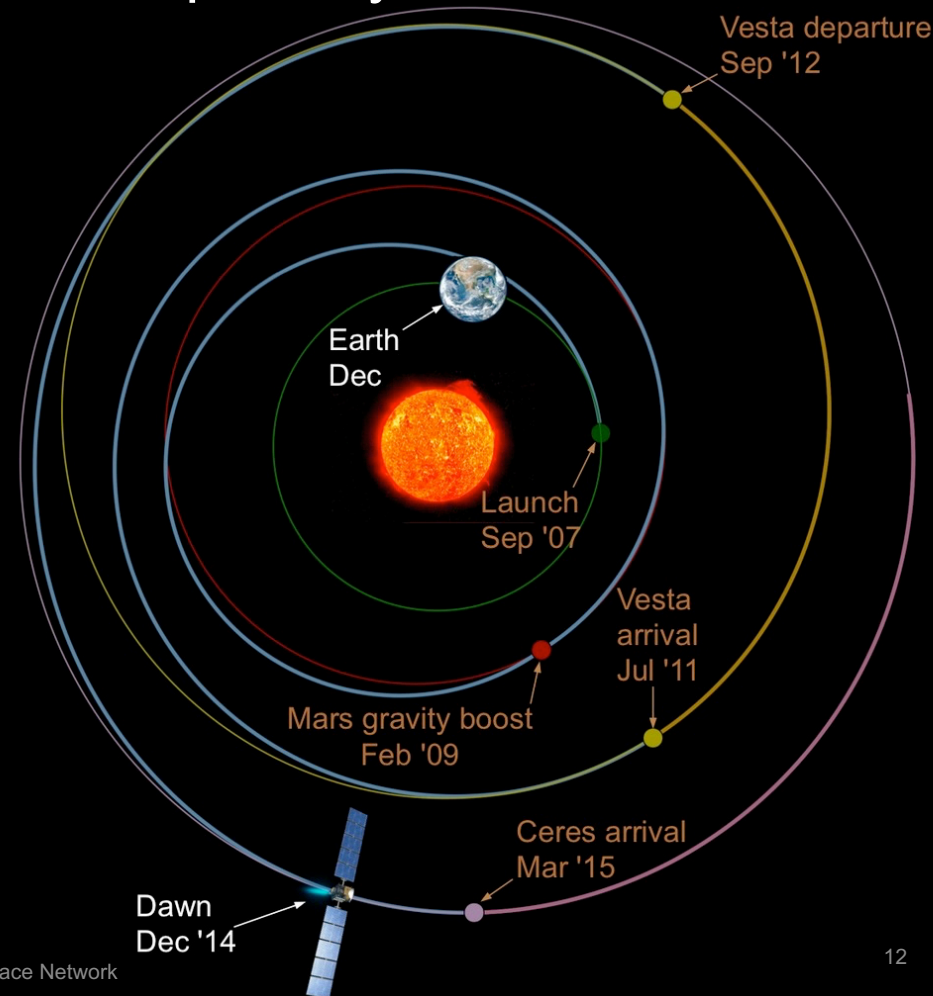
- There is no GPS in deep space
- Measurements of radio signal are the primary observables

Ranging: measurement of the distance to the spacecraft

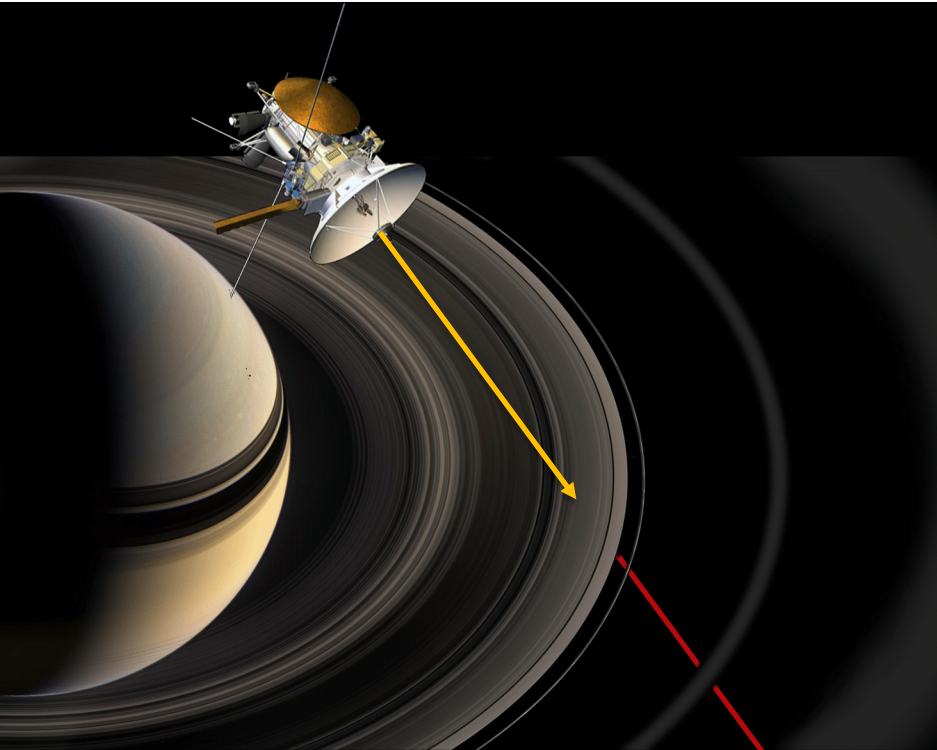
Doppler: measurement of the relative spacecraft motion

Delta Differenced One-Way Ranging (Δ DOR): Using multiple ground antennas to measure angle on the sky

- Supplemented with on-board sensors



DSN Science



- Measuring perturbations in the link

Attenuations

Spacecraft wobble

Frequency deviation

- We learn things about

Rings and particles

Atmospheres

Interiors of bodies

- We even use the DSN as a radar

See through atmospheres

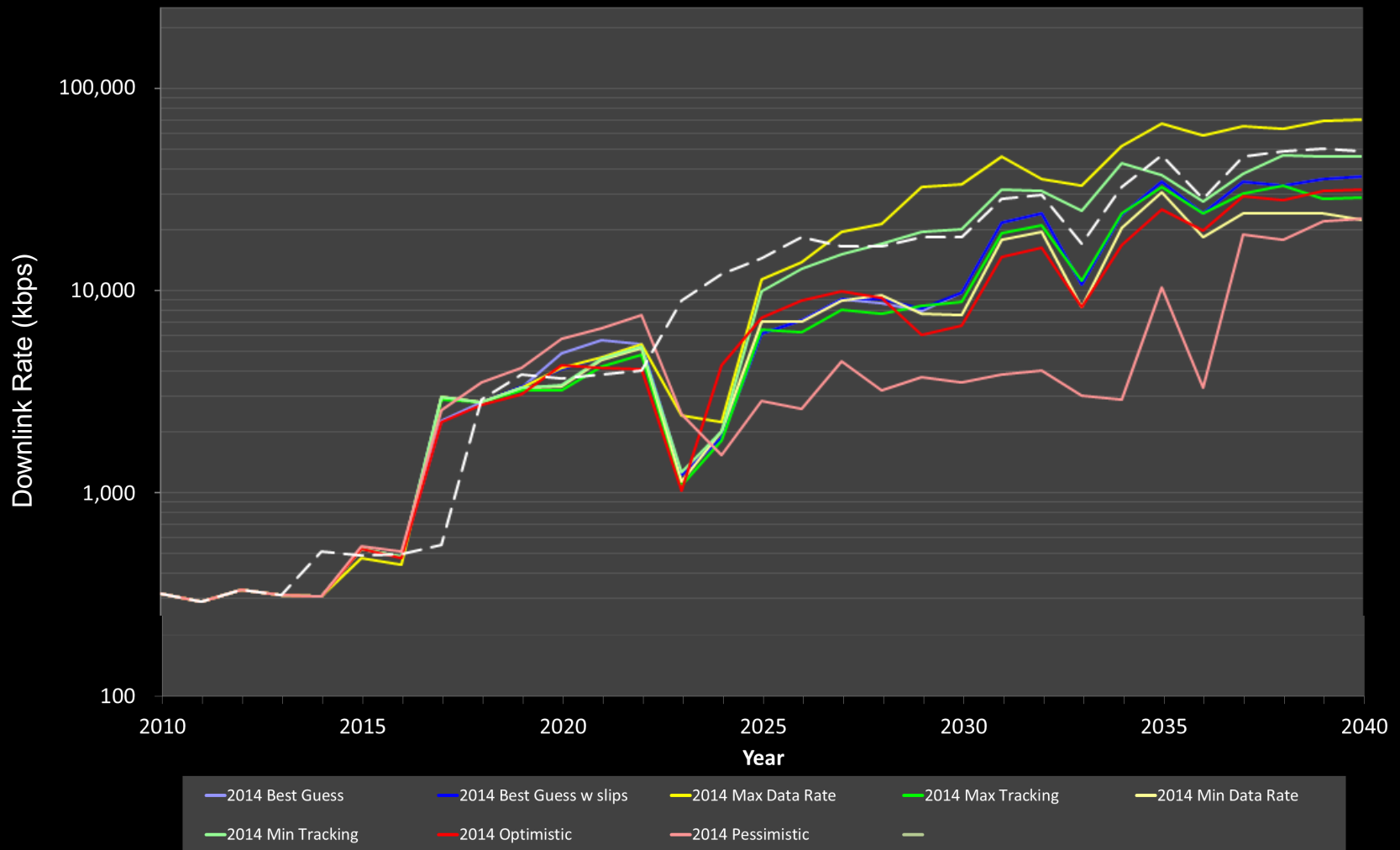
Study terrain

Assess danger from asteroids



Challenge: Future Missions Generate More Data

Average Across Each Mission's Maximum Downlink Rate as a Function of Time
(Comparison of Mission Set Scenarios)



Optical Communication in the DSN

- We will demonstrate deep space optical communications on the 2022 Psyche mission
- Uses Palomar 200" – but we need an operational capability after that
- Add mirrors to 34m DSN antennas to provide an equivalent 8m spherical aperture
- Place a photon-counting optical detector at apex
- Use separate, much smaller aperture for uplink, reducing requirements on this larger system

